

Prospects of Radio Astrometry with VLBI Networks

Phillip Galli (IAG/USP)

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São Paulo, Setembro de 2017.*



Astrometry

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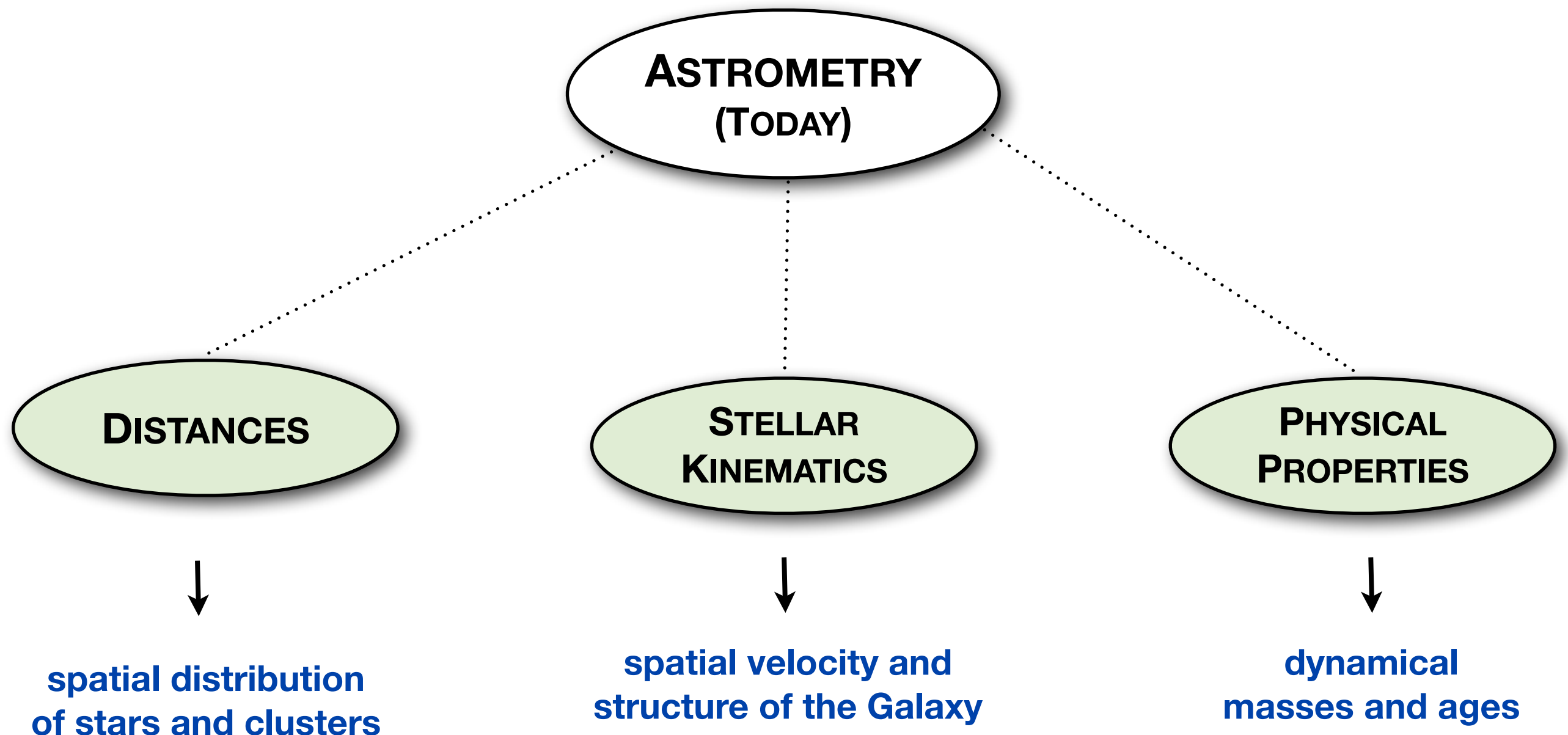
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OF THE CELESTIAL OBJECTS.**

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SYSTEMS, COORDINATE SYSTEMS, ETC.**

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- ✓ MICRO-ARCSECOND ASTROMETRY
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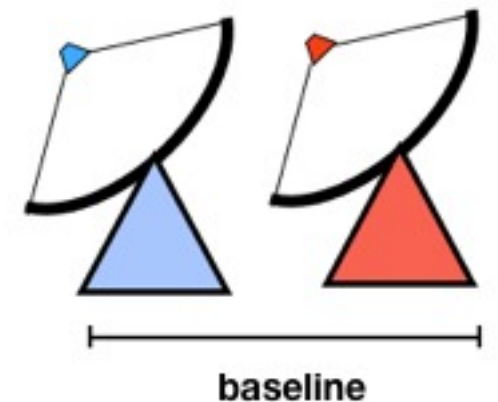
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VERY LONG BASELINE
INTERFEROMETRY (VLBI)



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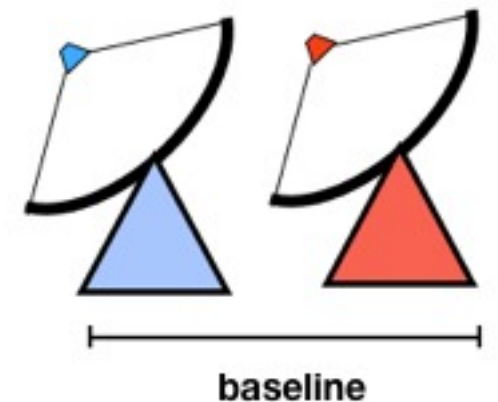
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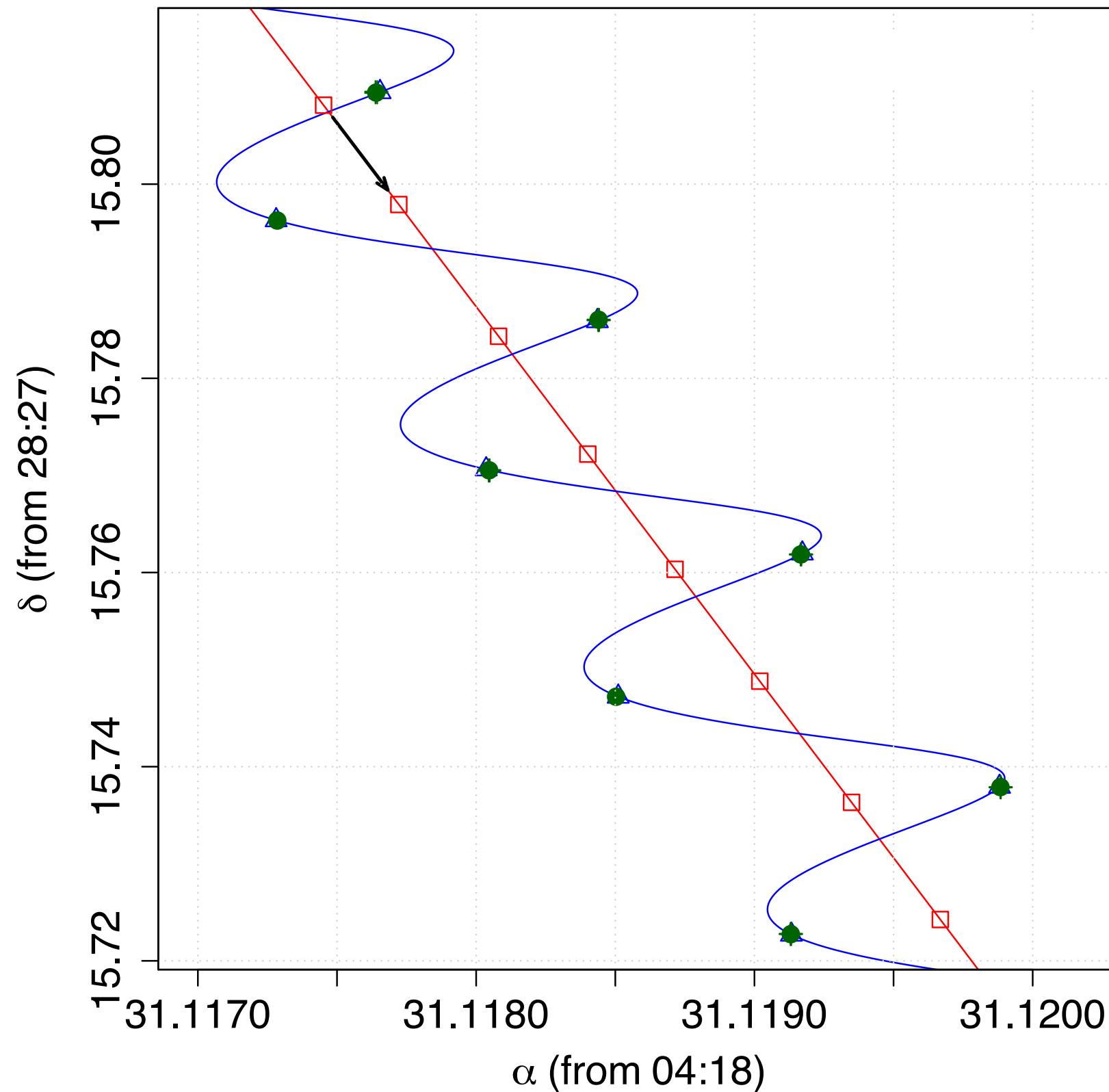
- ✓ MICRO-ARCSECOND ASTROMETRY
- ✓ RADIO: HIGH EXTINCTION
- ✓ TARGET-ORIENTED (> 5 YEARS)

The Gould Belt's Distances Survey (GOBELINS)

- **Aims:** stellar distances and proper motions of young stars (with unprecedented accuracy).
- **Targets:** nearby star-forming regions (**Taurus**, Ophiuchus, Orion, Serpens and Perseus).
- **Methods:** multi-epoch VLBI astrometry.
- **Instrument:** Very Long Baseline Array (VLBA)
 - * transcontinental baselines (~ 8000 km)
 - * Angular resolution: ~1 mas
 - * Astrometry precision: 10-50 μ as
- **Project:**
 - * 2000 hours with the VLBA (PI: L. Loinard)
 - * start: 2012, end: 2017/2018



Example 1: V410Tau (HD 283518)



This Work

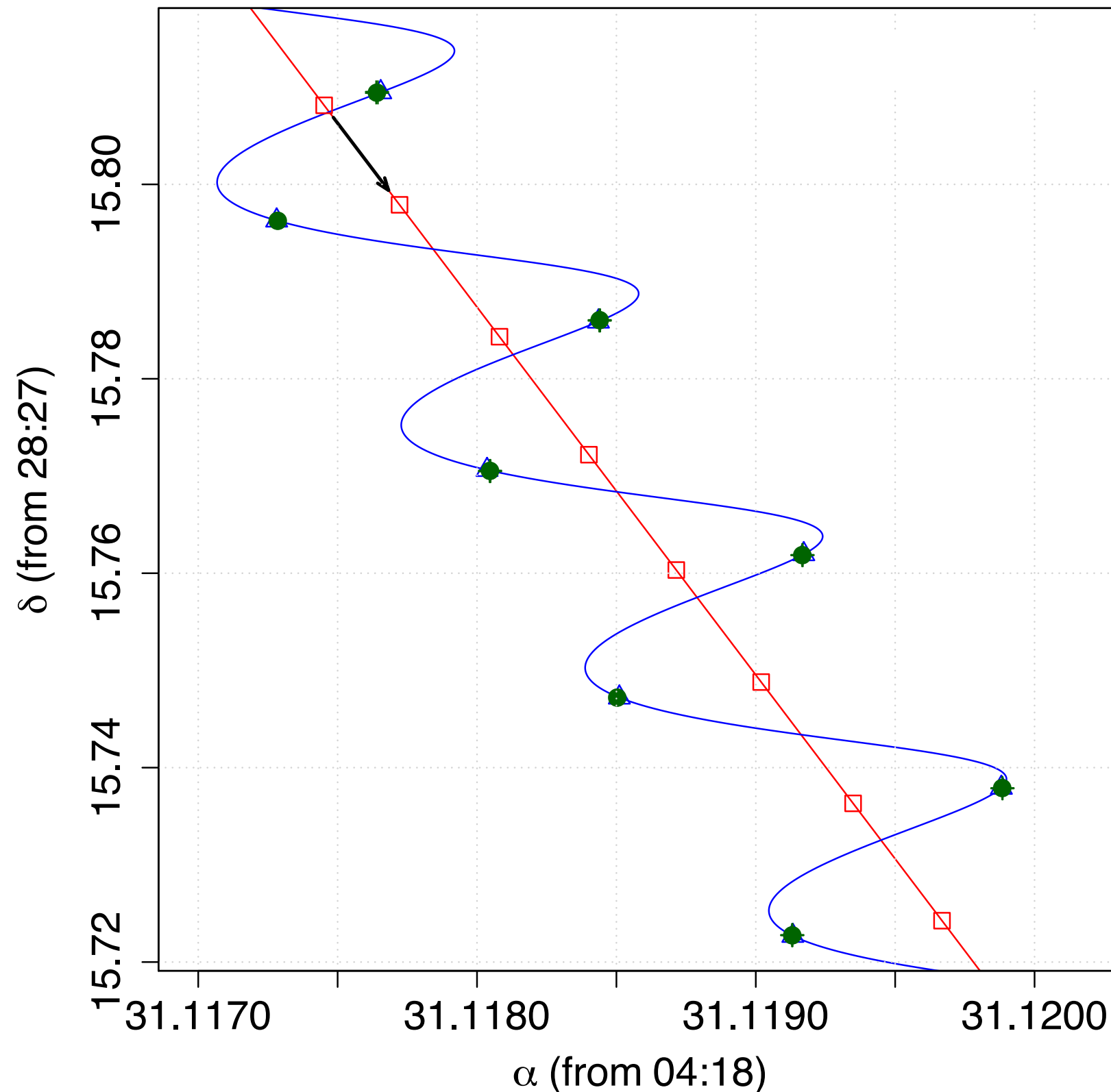
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$$\pi = 7.755 \pm 0.038 \text{ mas}$$

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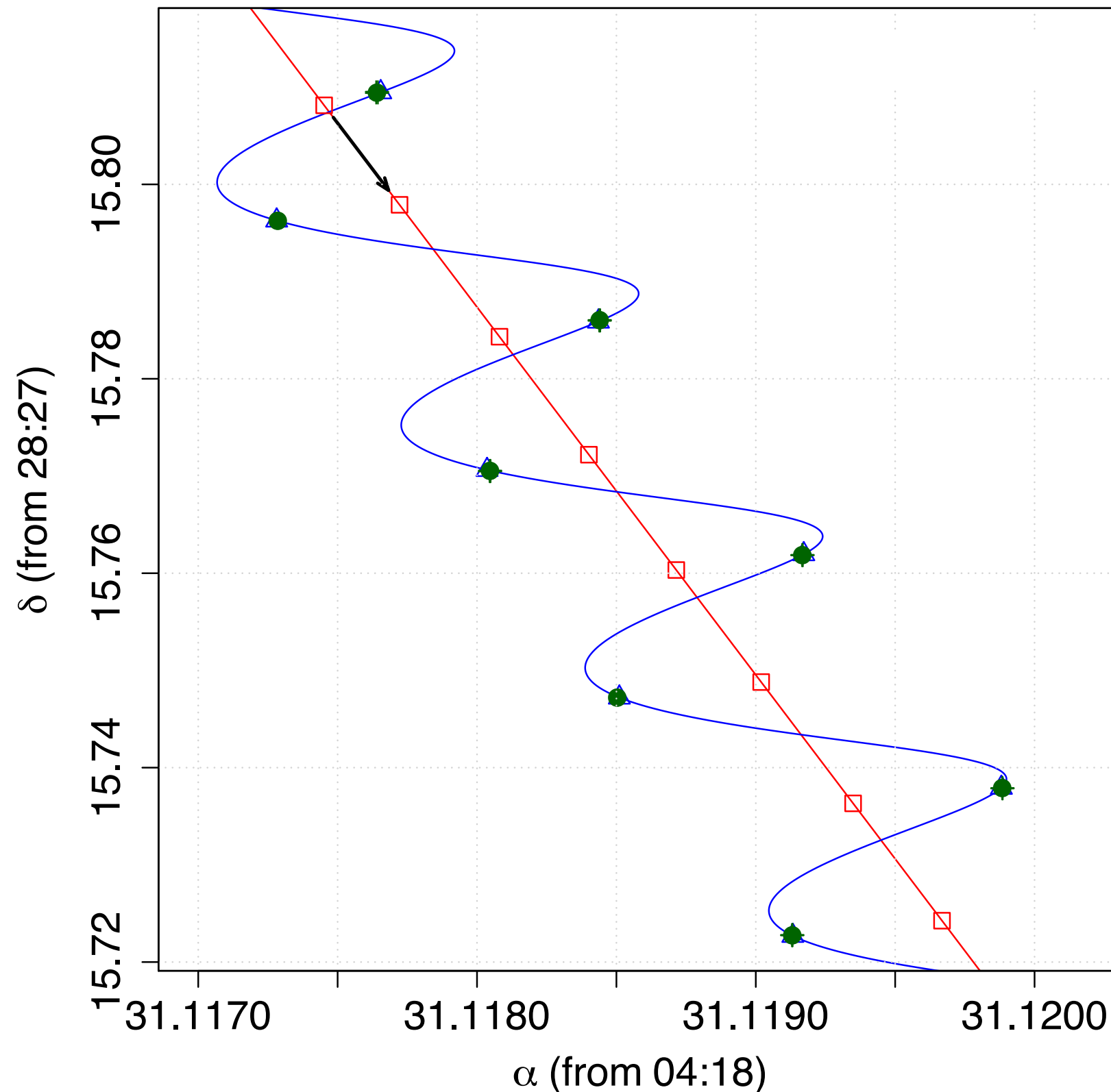
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Hipparcos

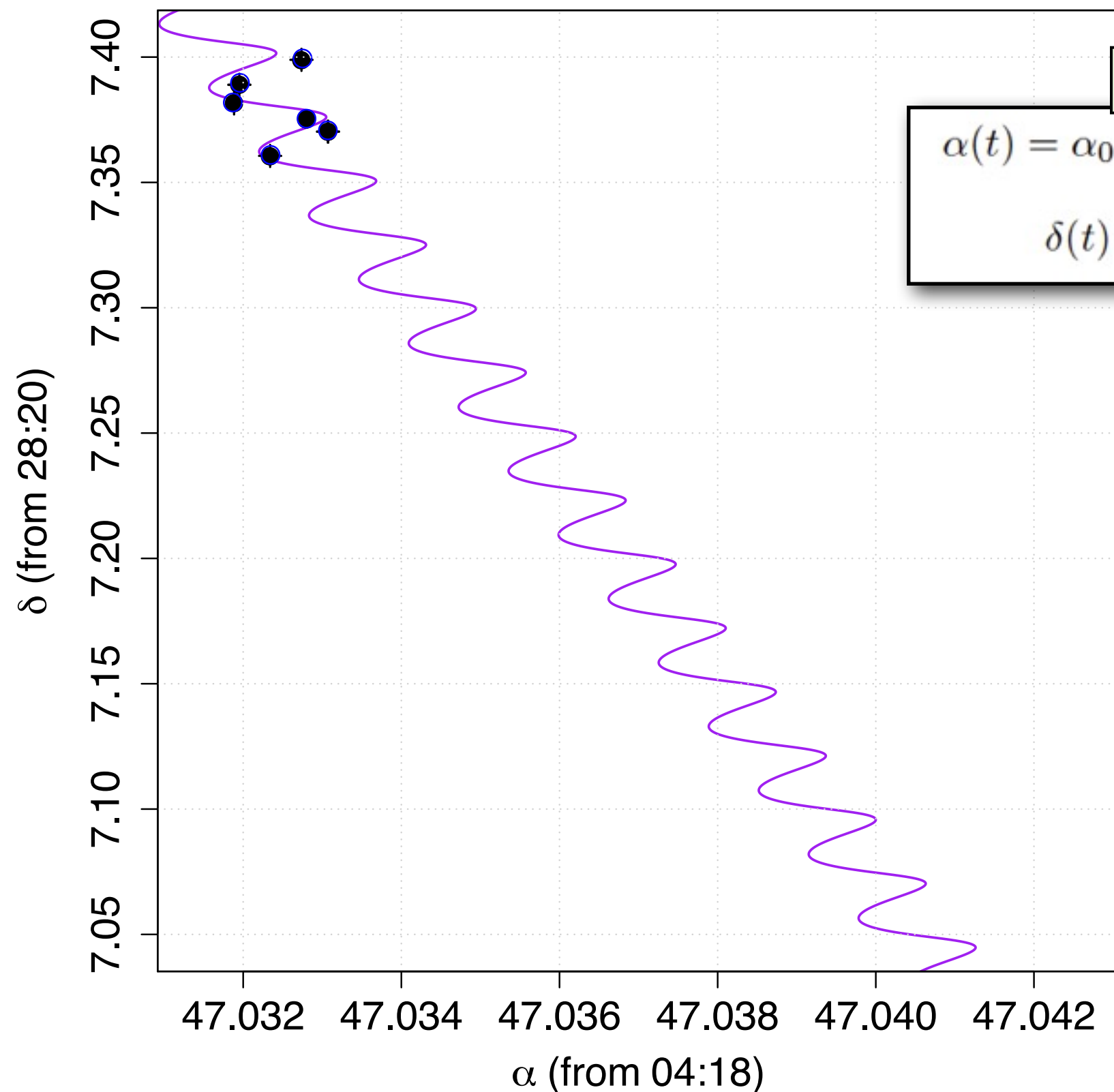
$$\mu_{\alpha}\cos\delta = 7.28 \pm 2.75 \text{ mas/yr}$$

$$\mu_{\delta} = -27.77 \pm 2.03 \text{ mas/yr}$$

$$\pi = 10.18 \pm 2.40 \text{ mas}$$

$$d = 98 [+30, -19] \text{ pc}$$

Example 2: V1023Tau (Hubble 4)

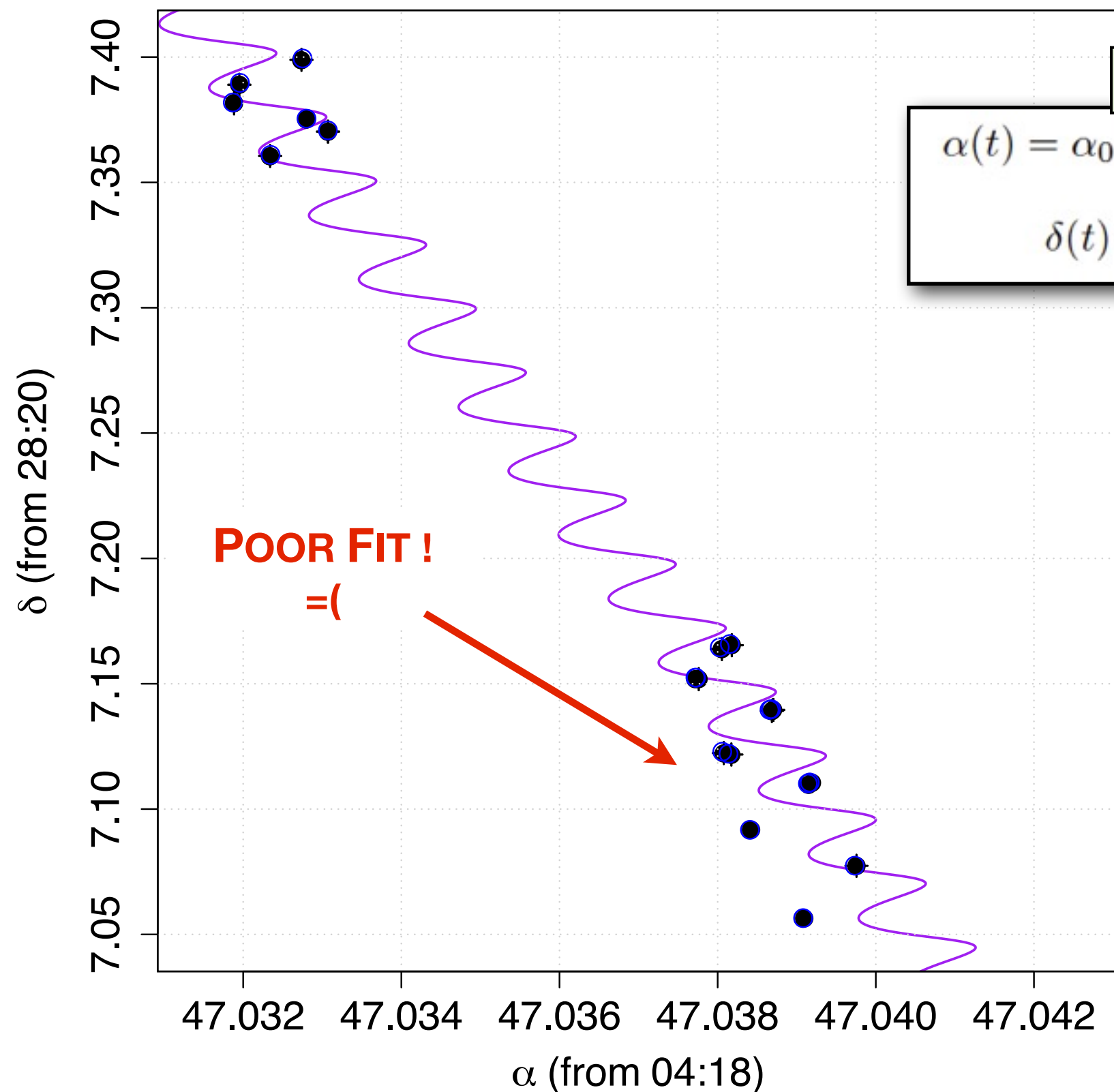


SYSTEM OF EQUATIONS

$$\alpha(t) = \alpha_0 + (\mu_\alpha \cos \delta)t + \pi f_\alpha(t)$$

$$\delta(t) = \delta_0 + \mu_\delta t + \pi f_\delta(t)$$

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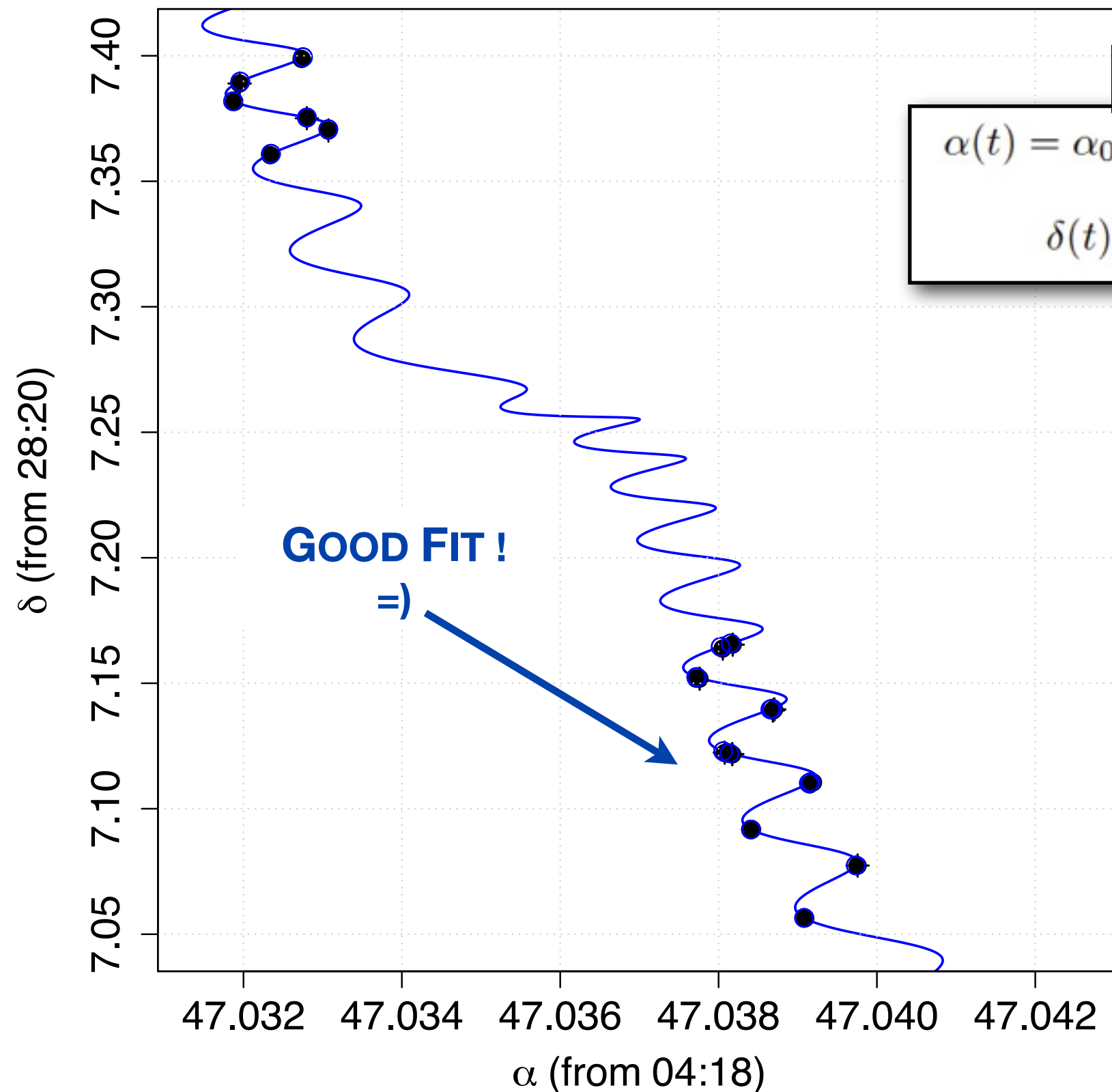


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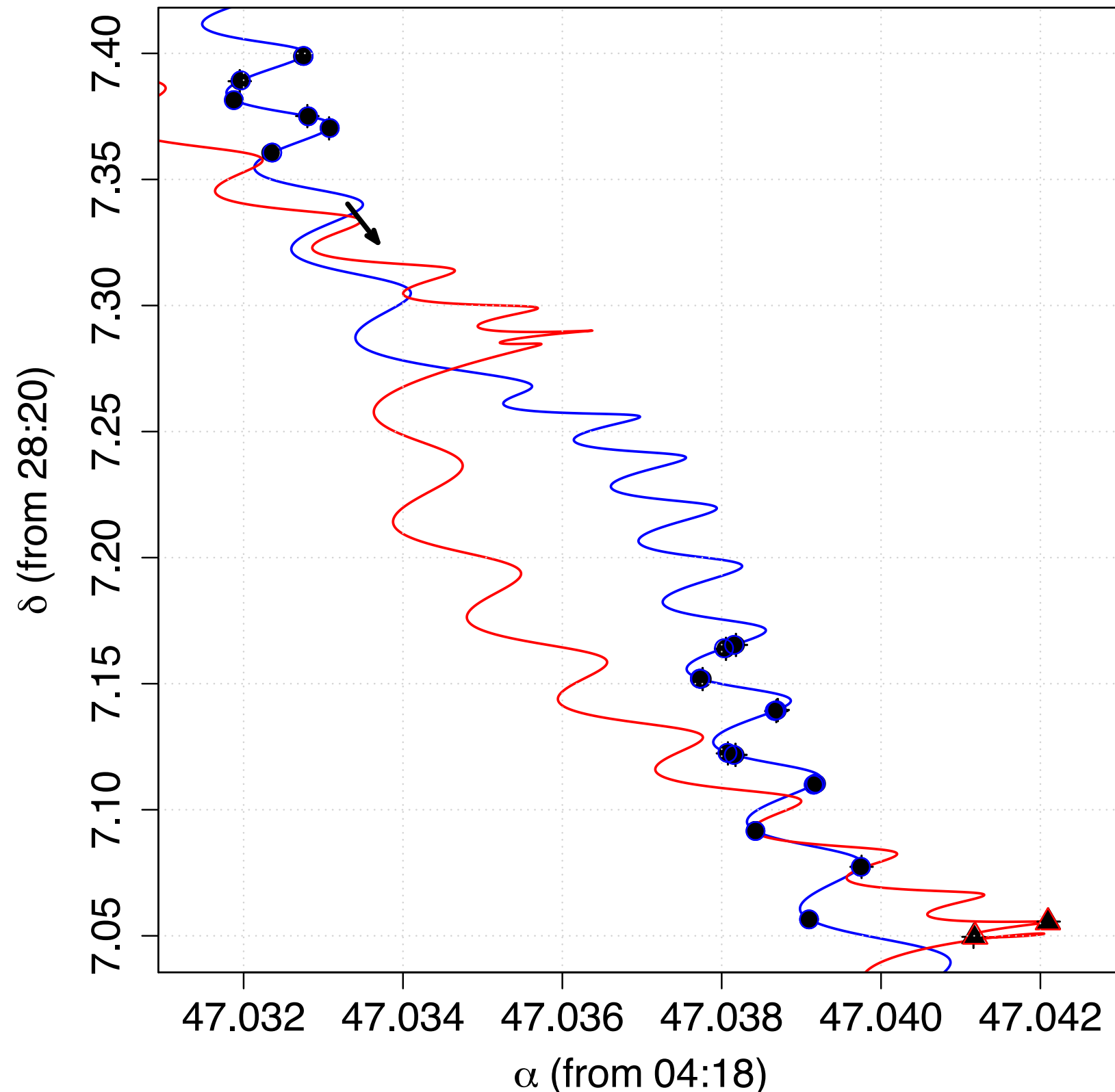
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ORBITAL MOTION !!!

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ASTROMETRY

$$\begin{aligned}\mu_{\alpha}\cos\delta &= 10.836 \pm 0.013 \text{ mas/yr} \\ \mu_{\delta} &= -25.448 \pm 0.014 \text{ mas/yr} \\ \pi &= 7.589 \pm 0.054 \text{ mas} \\ d &= 131.8 \pm 0.9 \text{ pc}\end{aligned}$$

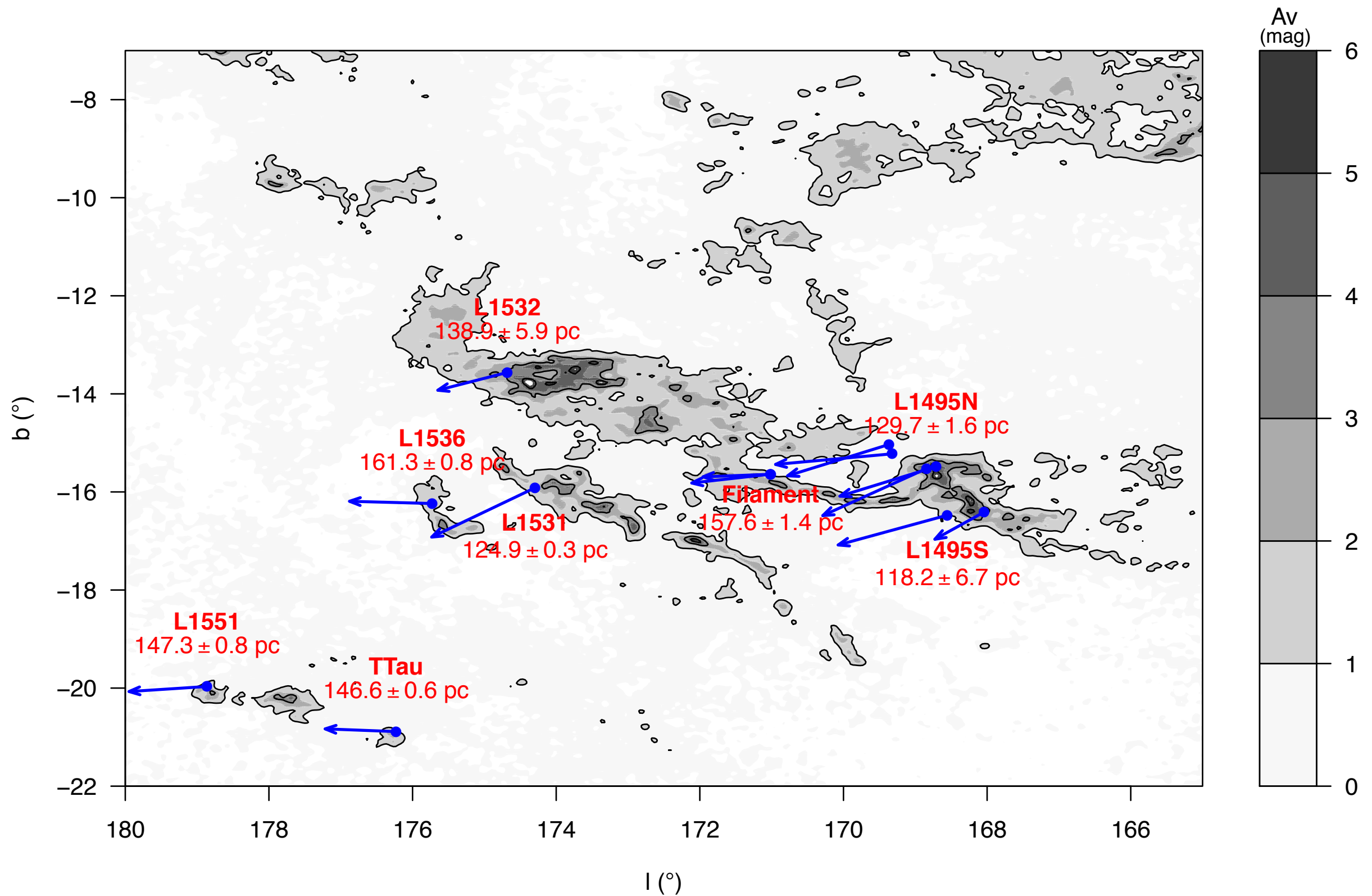
ORBITAL MOTION

$$\begin{aligned}a_1 &= 2.411 \pm 0.023 \text{ au} \\ a_2 &= 4.329 \pm 0.170 \text{ au} \\ P &= 9.171 \pm 0.010 \text{ yr} \\ e &= 0.767 \pm 0.001 \\ T_P &= 2458035.7 \pm 6.2 \text{ JD} \\ \Omega &= 70.1^\circ \pm 4.3^\circ \\ \omega &= 233.7^\circ \pm 3.9^\circ \\ i &= 40.8^\circ \pm 1.0^\circ\end{aligned}$$

DYNAMICAL MASSES

$$\begin{aligned}m_1 &= 2.38 \pm 0.11 M_{\odot} \\ m_2 &= 1.33 \pm 0.13 M_{\odot}\end{aligned}$$

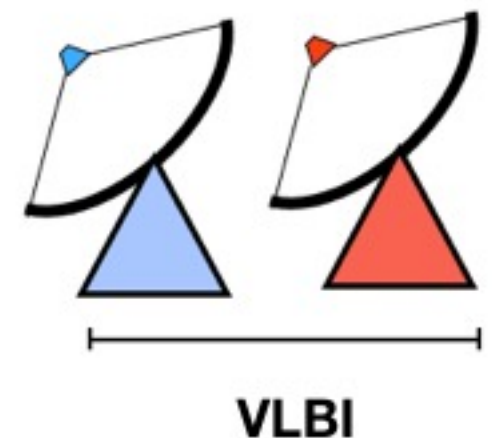
Distance to the Taurus Star-Forming Region



Conclusions

1) RADIO ASTROMETRY:

- * VLBI observations can deliver state-of-the-art astrometry.
- * complement the Gaia satellite from the ground.



2) THE LLAMA PROJECT AND SPANET:

- * construct the first Latin American VLBI network (targets in the southern sky)
- * develop a new research field (“radio astrometry”) in Brazil.

